

1 **Atmospheric dynamics feedback: concept, simulations and climate**
2 **implications**

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ABSTRACT

10 The regional climate response to radiative forcing is largely controlled by
11 changes in the atmospheric circulation. It has been suggested that global cli-
12 mate sensitivity also depends on the circulation response, an effect we call the
13 “atmospheric dynamics feedback”. Using a technique to isolate the influence
14 of changes in atmospheric circulation on top-of-atmosphere radiation, we cal-
15 culate the atmospheric dynamics feedback in coupled climate models. Large-
16 scale circulation changes contribute substantially to all-sky and cloud feed-
17 backs in the tropics but are relatively less important at higher latitudes. Glob-
18 ally averaged, the atmospheric dynamics feedback is positive and amplifies
19 the near-surface temperature response to climate change by an average of 8%
20 in simulations with coupled models. A constraint related to the atmospheric
21 mass budget results in the dynamics feedback being small on large scales rela-
22 tive to feedbacks associated with thermodynamic processes. Idealized-forcing
23 simulations suggest that circulation changes at high latitudes are potentially
24 more effective at influencing global temperature than circulation changes at
25 low latitudes, and the implications for past and future climate change are dis-
26 cussed.

27 **1. Introduction**

28 The general circulation of the atmosphere is expected to change in a variety of ways as climate
29 warms in response to CO₂-induced radiative forcing. Projected circulation responses include a
30 widening of the Hadley circulation (Lu et al. 2007), a narrowing of the intertropical convergence
31 zone (ITCZ) (Byrne and Schneider 2016b), a slowdown in the Walker circulation (Vecchi and So-
32 den 2007), and a poleward shift of the mid-latitude storm tracks (Yin 2005; Schneider et al. 2010).
33 Regional climate change, including projected changes in the water cycle, is strongly dependent
34 on these circulation responses (e.g., Seager et al. 2010; Scheff and Frierson 2012; Shepherd 2014;
35 Byrne and O’Gorman 2015). Not only does the circulation affect *regional* climate but it has also
36 been suggested that *global* climate change may be amplified or dampened by the atmospheric cir-
37 culation response via its effect on the planetary radiation budget (e.g., Pierrehumbert 1995); we
38 refer to this effect as the “atmospheric dynamics feedback”.

39 Studies of how atmospheric circulation responses to global warming affect top-of-atmosphere
40 (TOA) radiation and climate have largely focused on interactions between the tropical circulation
41 and clouds (Lindzen et al. 2001; Bony et al. 2004; Wyant et al. 2006; Mauritsen and Stevens 2015;
42 Bony et al. 2016). The ‘iris’ hypothesis of Lindzen et al. (2001) argued that a decrease in the
43 portion of the tropics covered by high clouds and a concurrent increase in low-cloud area under
44 global warming is an important negative feedback on climate change. Pierrehumbert (1995) put
45 forward a similar mechanism as a potential means of regulating tropical sea-surface temperatures
46 in warm climates of the past (e.g., the Eocene). The physics of this proposed negative feedback
47 are intuitive and rely on differences in cloud-radiative effect between low-cloud and high-cloud
48 regions. The shortwave and longwave cloud-radiative effects are defined conventionally as the
49 differences between the clear-sky outgoing shortwave or longwave fluxes at TOA and the all-

50 sky outgoing fluxes (Allan 2011). Specifically, the shortwave cloud-radiative effect is $SW_{\text{clear}}^{\text{TOA}} \uparrow$
51 $-SW_{\text{all}}^{\text{TOA}} \uparrow$, and the longwave cloud-radiative effect is $LW_{\text{clear}}^{\text{TOA}} \uparrow - LW_{\text{all}}^{\text{TOA}} \uparrow$, where SW and LW
52 indicate the longwave and shortwave TOA fluxes, respectively, and \uparrow denotes an upward flux. Low
53 clouds have a negative cloud-radiative effect and cool the Earth because they reflect incoming
54 shortwave radiation while having only a weak longwave greenhouse effect. High clouds, on the
55 other hand, typically have a small cloud-radiative effect because of a large cancellation between
56 shortwave cooling and longwave warming (Ramanathan et al. 1989). Consequently, a decrease in
57 the area of high clouds (with small cloud-radiative effect) and an increase in the area of strongly
58 cooling low clouds under global warming is a possible negative feedback on climate change.

59 Although the validity of the original iris hypothesis proposed by Lindzen et al. (2001) has been
60 strongly challenged (e.g., Fu et al. 2002), recent studies have revived interest in iris-type phe-
61 nomena as potentially important feedbacks on climate change (Mauritsen and Stevens 2015; Bony
62 et al. 2016; Cronin and Wing 2017). The suggestion that tropical convection will become more
63 aggregated as sea-surface temperatures increase (e.g., Wing and Emanuel 2014) has formed the
64 conceptual basis for these ‘iris 2.0’ studies. Convective aggregation is not typically simulated
65 by global climate models, in which convection is parameterized out of necessity, but it has been
66 put forward as a negative iris-type atmospheric dynamics feedback (Bony et al. 2016; Wing et al.
67 2017; Cronin and Wing 2017). Perhaps related to convective aggregation, a narrowing of trop-
68 ical high-cloud regions (i.e., the ITCZ) relative to neighboring low-cloud regions under global
69 warming has been observed over recent decades (Wodzicki and Rapp 2016) and further narrowing
70 is expected as the climate continues to warm (Lau and Kim 2015; Byrne and Schneider 2016b).
71 This ITCZ narrowing represents another potential iris effect. At higher latitudes, cloud feedbacks
72 associated with shifts in the extratropical circulation have received considerable attention in re-

73 cent years (Grise et al. 2013; Kay et al. 2014; Wall and Hartmann 2015), though climate models
74 disagree on the magnitude and importance of these feedbacks (Ceppi et al. 2017a).

75 The studies mentioned above focus on cloud feedbacks, but it is plausible that a re-organization
76 of the atmospheric circulation could also affect the water-vapor feedback by changing the spatial
77 pattern and intensities of moisture convergence and divergence zones. More generally, atmo-
78 spheric circulation changes can conceivably influence both the cloud-radiative effect and all-sky
79 TOA radiative fluxes at all latitudes, but the magnitudes of these atmospheric dynamics feedbacks
80 and their dependences on latitude are unknown. Are there latitudes at which climate feedbacks
81 are dominated by large-scale circulation changes? How do these circulation changes contribute to
82 global climate change? Can our emerging understanding of atmospheric dynamics in a changing
83 climate be used to better constrain global climate sensitivity? Are low-latitude or high-latitude cir-
84 culation changes more effective at changing global-mean surface temperature? Related to this lat-
85 ter question, a number of studies have shown that the spatial pattern of ocean heat uptake strongly
86 affects the global temperature response (Armour et al. 2013; Rose et al. 2014; Rugenstein et al.
87 2016), i.e., heat absorbed by the ocean at high latitudes has a stronger effect on global temperature
88 than the same quantity of heat absorbed at low latitudes. Should such a latitudinal dependence also
89 exist for the atmospheric dynamics feedback, it would have important implications for the ability
90 of tropical iris-type mechanisms to regulate past and future climate change.

91 Here we use coupled climate models together with idealized simulations to assess how large-
92 scale atmospheric circulation changes contribute to TOA radiative feedbacks and global climate
93 change. We begin by outlining the method used to decompose dynamic versus thermodynamic
94 influences on TOA radiation (Section 2) before analysing the atmospheric dynamics feedback in
95 simulations with coupled models (Section 3). Idealized simulations with stylized TOA radiative
96 forcings are also performed to understand how circulation changes at different latitudes could

97 influence global climate (Section 4). We conclude by summarizing our results and discussing
98 directions for future research (Section 5).

99 **2. Estimating the atmospheric dynamics feedback**

100 *a. CMIP5 simulations*

101 We will investigate the role of the atmospheric circulation in controlling radiative feedbacks
102 using simulations from the Coupled Model Intercomparison Project Phase 5 (CMIP5) (Taylor
103 et al. 2012). In particular, we analyze changes in TOA radiative fluxes for 27 models¹ be-
104 tween pre-industrial control runs (*piControl*) and runs in which the CO₂ concentration is instan-
105 taneously quadrupled relative to pre-industrial levels (*abrupt4xCO2*). For both the *piControl* and
106 *abrupt4xCO2* runs, we analyze 100 years of monthly-mean data. All-sky and clear-sky TOA ra-
107 diative fluxes are defined as positive downwards.

108 *b. Dependence of top-of-atmosphere radiation on the atmospheric circulation*

109 To estimate the radiative feedbacks associated with changes in the large-scale atmospheric cir-
110 culation, we first decompose TOA radiative anomalies (e.g., due to changes in greenhouse gas
111 concentrations) into “dynamic”, “thermodynamic”, and “nonlinear” components (where nonlin-
112 ear here means the combined influence of changes in dynamic and thermodynamic processes on
113 TOA radiation). To perform this decomposition, we follow Bony et al. (2004) in using the mid-
114 tropospheric vertical velocity, ω , as an admittedly imperfect proxy for dynamical effects. That

¹The models analyzed in this study are: ACCESS1-0, ACCESS1-3, BCC-CSM1-1, BCC-CSM1-1-M, CanESM2, CCSM4, CNRM-CM5, CNRM-CM5-2, CSIRO-Mk3-6-0, FGOALS-g2, FGOALS-s2, GFDL-CM3, GFDL-ESM2G, GFDL-ESM2M, GISS-E2-H, GISS-E2-R, HadGEM2-ES, INMCM4, IPSL-CM5A-LR, IPSL-CM5B-LR, MIROC5, MIROC-ESM, MPI-ESM-LR, MPI-ESM-MR, MPI-ESM-P, MRI-CGCM3, and NorESM1-M

115 is, we decompose changes in TOA fluxes into those that are, and are not, associated with vertical
 116 velocity changes.

117 To estimate how changes in atmospheric circulation affect the TOA radiative response to an im-
 118 posed forcing, we first construct discretized functions $R_{\text{cloud}}(\omega, \phi)$ and $R_{\text{all}}(\omega, \phi)$ at each latitude
 119 ϕ , where $R_{\text{cloud}}(\omega, \phi)$ is the cloud-radiative effect and $R_{\text{all}}(\omega, \phi)$ is the all-sky TOA radiative flux,
 120 and both quantities are ‘binned’ as a function of mid-tropospheric vertical velocity. The radiative
 121 fluxes and vertical velocities used in our analysis are all monthly-mean quantities. We also con-
 122 struct normalized area probability density functions (PDFs) for vertical velocity at each latitude,
 123 $A(\omega, \phi)$. Note that $R_{\text{cloud}}(\omega, \phi)$, $R_{\text{all}}(\omega, \phi)$, and $A(\omega, \phi)$ (and the radiative anomalies and feed-
 124 backs discussed below) are calculated for each CMIP5 model individually before being linearly
 125 interpolated to a common grid for the purpose of plotting multimodel means as a function of lati-
 126 tude. In this framework, the zonally-averaged cloud-radiative effect or all-sky TOA flux [$\overline{R_{\text{cloud}}}(\phi)$
 127 or $\overline{R_{\text{all}}}(\phi)$] is the product of $R_{\text{cloud}}(\omega, \phi)$ or $R_{\text{all}}(\omega, \phi)$ with $A(\omega, \phi)$ at latitude ϕ integrated over
 128 all vertical-velocity bins, which in continuous form are expressed as:

$$\overline{R_{\text{cloud}}}(\phi) = \int_{-\infty}^{\infty} R_{\text{cloud}}(\omega, \phi) A(\omega, \phi) d\omega \quad (1)$$

$$\overline{R_{\text{all}}}(\phi) = \int_{-\infty}^{\infty} R_{\text{all}}(\omega, \phi) A(\omega, \phi) d\omega, \quad (2)$$

129 where $\overline{(\cdot)}$ denotes an area-weighted integral over vertical velocity. The $R_{\text{cloud}}(\omega, \phi)$, $R_{\text{all}}(\omega, \phi)$,
 130 and $A(\omega, \phi)$ functions for the *piControl* runs are constructed for each climatological month us-
 131 ing 100 years of monthly-mean simulation data. These functions are then re-calculated for each
 132 month and year in the *abrupt4xCO2* runs to quantify [using the decompositions (4) and (5) de-
 133 scribed below] how dynamic and thermodynamic processes contribute to the radiative feedbacks
 134 that determine Earth’s response to a quadrupling of CO₂.

135 The dependence of tropical cloud-radiative effect on vertical velocity (Fig. 1a) and the area PDF
136 of vertical velocity (Fig. 1d) in one coupled climate model are similar to the observed functions
137 [cf. Fig. 2 in Bony et al. (2004)]. Note that in Figure 1, each quantity has been averaged from
138 $\phi = -30^\circ$ to 30° with area weighting. Between ascending ($\omega < 0$) and weakly descending ($\omega > 0$)
139 regions, the longwave and shortwave cloud-radiative effects vary approximately linearly with mid-
140 tropospheric vertical velocity (Fig. 1a). For the longwave cloud-radiative effect, the positive
141 (i.e., heating) values in strongly ascending regions such as the ITCZ reflect cold, high-altitude
142 cloud tops and large cloud fractions, which together produce a strong greenhouse effect. The
143 decreasing longwave cloud-radiative effect with increasing ω (i.e., weaker upward motion) is due
144 to cloud-top altitudes and cloud fractions declining for weaker ascent regimes. For the shortwave
145 cloud-radiative effect (Fig. 1a), the reduced cooling as ascending motion weakens is a result
146 of decreasing cloud fractions and less shortwave reflection to space. The relative insensitivity
147 of the longwave and shortwave cloud-radiative effects to changes in vertical velocity in strongly
148 descending regions is likely due to cloud-masking effects (Soden et al. 2004). It shows that the
149 altitude and fraction of low clouds in such regions no longer depend strongly on mid-tropospheric
150 vertical velocity (Bony et al. 2004).

151 Although we would not necessarily expect the dependence of clear-sky radiative fluxes on verti-
152 cal velocity to be as strong as that found for the cloud-radiative effect, it is plausible that regions of
153 low-level convergence would be moister than regions of divergence, which would have a signature
154 in the clear-sky radiative flux and contribute to an all-sky flux dependence on vertical velocity. In-
155 deed tropical ascending regions are observed to have smaller clear-sky longwave cooling relative
156 to descending regions (Bony et al. 1997); this clear-sky flux dependence on vertical velocity is at
157 least qualitatively captured by climate models (Fig. 1b). There is also a simulated dependence
158 of clear-sky shortwave radiation on vertical velocity (Fig. 1b), with regions of descent generally

159 having a larger outgoing shortwave flux and a stronger longwave radiative cooling (presumably
 160 because the atmosphere in these descending regions is dry and consequently absorbs less short-
 161 wave radiation, for example above deserts such as the Sahara). The influence of vertical velocity
 162 on clear-sky fluxes and on the cloud-radiative effect combine to produce a strong dependence of
 163 all-sky fluxes on vertical velocity, particularly the longwave fluxes (Fig. 1c). Figure 1 shows
 164 the relationships of cloud-radiative effect, clear-sky fluxes, and all-sky fluxes to vertical velocity
 165 for one climate model, but these relationships are qualitatively similar for the other 26 models
 166 analyzed (not shown).

167 The area PDF of vertical velocity in a climate model shows a skewed distribution in the tropics
 168 (Fig. 1d), with a large area covered by slow-moving descending air masses — the so-called ‘ra-
 169 diator fin’ regions where longwave cooling is strong (Pierrehumbert 1995). The area PDF has a
 170 relatively small ascent region with a larger variance in vertical velocity compared to the descent
 171 region (Fig. 1d), and is qualitatively similar to a PDF calculated using reanalysis data (Bony et al.
 172 2004).

173 *c. Dynamic/thermodynamic decomposition of changes in top-of-atmosphere radiation*

174 Our objective is to identify the roles of dynamic versus thermodynamic processes in control-
 175 ling the *response* of TOA radiation to an imposed $4 \times \text{CO}_2$ forcing, i.e., we want to estimate the
 176 atmospheric dynamics and thermodynamics feedbacks. Consequently, before applying the de-
 177 composition method of Bony et al. (2004), we first need to estimate and remove the component
 178 of the TOA radiative anomalies associated with the imposed forcing. To do this, we consider a
 179 linearization of the TOA radiation budget which is typically invoked to describe global climate
 180 change (e.g., Armour et al. 2013):

$$[\delta R_{\text{all}}]^g = [F]^g + [\lambda_{\text{all}}]^g [\delta T]^g, \quad (3)$$

181 where δR_{all} denotes the all-sky TOA radiative imbalance (or anomaly) in units of W m^{-2} , F is
 182 the imposed radiative forcing (W m^{-2}), λ_{all} is the all-sky feedback parameter ($\text{W m}^{-2} \text{K}^{-1}$), δT
 183 is the surface-air temperature response (K), δ indicates an anomaly (or change) between the *pi-*
 184 *Control* and *abrupt4xCO2* runs, and $[\cdot]^g$ denotes a global average. The all-sky feedback parameter
 185 quantifies the efficiency with which the Earth system equilibrates a radiative forcing. In this linear
 186 framework, the radiative response of the climate system to the imposed forcing is obtained by
 187 subtracting the forcing from the total TOA radiative anomaly, i.e., $[\delta R'_{\text{all}}]^g \equiv [\delta R_{\text{all}}]^g - [F]^g$ where
 188 $\delta R'_{\text{all}}$ is defined to be the *all-sky radiative response*.

189 The dynamic components of the (i) all-sky radiative response or (ii) changes in cloud-radiative
 190 effect are identified as the TOA radiative anomalies induced by changes in $A(\omega, \phi)$ for (i) fixed
 191 $R_{\text{all}}(\omega, \phi)$ or (ii) fixed $R_{\text{cloud}}(\omega, \phi)$. The decompositions are expressed mathematically as:

$$\begin{aligned} \overline{\delta R'_{\text{all}}}(\phi) &= \underbrace{\int_{-\infty}^{\infty} R_{\text{all}}(\omega, \phi) \delta A(\omega, \phi) d\omega}_{\text{dynamic}} + \underbrace{\int_{-\infty}^{\infty} [\delta R_{\text{all}}(\omega, \phi) - F(\phi)] A(\omega, \phi) d\omega}_{\text{thermodynamic}} \\ &+ \underbrace{\int_{-\infty}^{\infty} [\delta R_{\text{all}}(\omega, \phi) - F(\phi)] \delta A(\omega, \phi) d\omega}_{\text{nonlinear}} \end{aligned} \quad (4)$$

$$\begin{aligned} \overline{\delta R_{\text{cloud}}}(\phi) &= \underbrace{\int_{-\infty}^{\infty} R_{\text{cloud}}(\omega, \phi) \delta A(\omega, \phi) d\omega}_{\text{dynamic}} + \underbrace{\int_{-\infty}^{\infty} \delta R_{\text{cloud}}(\omega, \phi) A(\omega, \phi) d\omega}_{\text{thermodynamic}} \\ &+ \underbrace{\int_{-\infty}^{\infty} \delta R_{\text{cloud}}(\omega, \phi) \delta A(\omega, \phi) d\omega}_{\text{nonlinear}} \end{aligned} \quad (5)$$

192 where $\overline{\delta R'_{\text{all}}}(\phi)$ is the total (shortwave + longwave) all-sky radiative response at latitude ϕ averaged
 193 over all vertical velocities with area weighting (note that the forcing is removed on the right-
 194 hand side), $\overline{\delta R_{\text{cloud}}}(\phi)$ is the change in cloud-radiative effect, and $F(\phi)$ is the diagnosed radiative
 195 forcing (see Section 2d for details regarding the calculation of the forcing), and $\delta A(\omega, \phi)$ is the
 196 change in the normalized area PDF of vertical velocity. The anomalies $\overline{\delta R'_{\text{all}}}(\phi)$ and $\overline{\delta R_{\text{cloud}}}(\phi)$

197 are calculated by evaluating the changes in radiative fluxes in each of the first 100 years of the
198 *abrupt4xCO2* simulations relative to the climatological fluxes from the *piControl* simulations. By
199 construction, the decompositions (4) and (5) match the simulated all-sky radiative responses and
200 changes in cloud-radiative effects very accurately (not shown).

201 In general, the dynamic components of (4) and (5) could be due to changes in the areas occupied
202 by ascending versus descending air (e.g., a narrowing of the ITCZ), strengthening/weakening of
203 the circulations within the ascending and descending regimes (e.g., a slowdown of the Walker
204 circulation), or spatial shifts in circulation features (e.g., a poleward movement of the storm track).
205 It should be noted that dynamic processes, for example lower-tropospheric mixing or convective
206 mass fluxes, that might be de-coupled from mid-tropospheric vertical velocity or which are not
207 explicitly resolved by global climate models, do not form part of what we define as the (large-scale)
208 dynamic components and are instead included in the thermodynamic components (Wyant et al.
209 2006). The thermodynamic components represent the portion of the all-sky radiative response or
210 the change in cloud-radiative effect that would occur if the atmospheric circulation did not change
211 [i.e., for fixed $A(\omega, \phi)$], and includes effects such as changes in cloud optical depth and increases
212 in water-vapor concentration that accompany a warming atmosphere. The nonlinear component
213 represents the combined effect of changes in dynamic and thermodynamic processes. We apply
214 this decomposition not only to changes in the cloud-radiative effect as in Bony et al. (2004), but
215 also to the all-sky radiative response.

216 Our method of evaluating (4) and (5) differs from that of Bony et al. (2004) in a number of
217 respects. Firstly, we perform the decomposition over the whole globe, not only for tropical re-
218 gions, and construct the $R_{\text{all}}(\omega, \phi)$ and $R_{\text{cloud}}(\omega, \phi)$ functions and the area PDFs using a vertically-
219 averaged vertical velocity between 400 hPa and 500 hPa (our results are qualitatively similar when
220 we average vertical velocity over the full atmosphere). Secondly, to understand how circulation

221 changes at different latitudes contribute to the atmospheric dynamics feedback, we perform the
222 decomposition at each latitude individually rather than over a large region. If the cloud-radiative
223 effect and all-sky radiative flux were solely functions of vertical velocity, then it would be pos-
224 sible to construct R_{all} and R_{cloud} using global data and then trace back to latitude to understand
225 the influence of circulation changes in different regions on TOA radiation. However, TOA radia-
226 tion depends not only on vertical velocity but also on temperature, relative humidity, solar zenith
227 angle, surface albedo, and other factors. Consequently, building R_{all} and R_{cloud} functions using
228 global data and then using these functions to determine dynamics and thermodynamics feedbacks
229 at a particular latitude is not possible as the local TOA radiation budget does not close.

230 When performing the decompositions (4) and (5), the vertical-velocity limits of $R_{\text{all}}(\omega, \phi)$,
231 $R_{\text{cloud}}(\omega, \phi)$, and $A(\omega, \phi)$ are specified to be the maximum and minimum values of the vertically-
232 averaged vertical velocity at each latitude, and we use 21 equally-spaced vertical-velocity ‘bins’.
233 If there are bins that are empty when constructing $R_{\text{all}}(\omega, \phi)$, $R_{\text{cloud}}(\omega, \phi)$, and $A(\omega, \phi)$, we use
234 linearly interpolated forms of these functions for which the missing values are filled in when cal-
235 culating the various terms in (4) and (5). We now use these decompositions to estimate the climate
236 feedbacks associated with dynamic and thermodynamic processes.

237 *d. Feedback calculation*

238 We define the dynamics, thermodynamics, and nonlinear all-sky and cloud feedbacks at each
239 latitude to be the respective components of the all-sky radiative response (4) or changes in cloud-
240 radiative effect (5) at that latitude normalized by the global-mean surface-air temperature change
241 [feedbacks are conventionally defined in climate science as TOA radiative responses normalized

242 by the surface temperature change (e.g., Rose and Rayborn 2016)]:

$$\lambda_{\text{all}}(\phi) = \lambda_{\text{all}}^{\text{dy}}(\phi) + \lambda_{\text{all}}^{\text{th}}(\phi) + \lambda_{\text{all}}^{\text{nl}}(\phi) \quad (6)$$

$$\lambda_{\text{cloud}}(\phi) = \lambda_{\text{cloud}}^{\text{dy}}(\phi) + \lambda_{\text{cloud}}^{\text{th}}(\phi) + \lambda_{\text{cloud}}^{\text{nl}}(\phi), \quad (7)$$

243 where λ_{all} is the all-sky feedback, λ_{cloud} is the cloud feedback, and the labels “dy”, “th”, and “nl”
244 denote the dynamic, thermodynamic, and nonlinear components of each feedback, respectively.
245 Our focus in this study is the atmospheric dynamics feedback, which we define to be the $\lambda_{\text{all}}^{\text{dy}}(\phi)$
246 component; we will also discuss the dynamic component of the cloud feedback, $\lambda_{\text{cloud}}^{\text{dy}}(\phi)$. The
247 global average of the all-sky feedback is simply the global feedback parameter defined in (3). The
248 various components of the all-sky (6) and cloud feedbacks (7) are estimated for each climate model
249 individually using ordinary least-squares regressions of the dynamic, thermodynamic, and nonlin-
250 ear components of $\overline{\delta R'_{\text{all}}}(\phi)$ and $\overline{\delta R'_{\text{cloud}}}(\phi)$ versus global-mean surface-air temperature changes
251 following CO₂ quadrupling (all anomalies in these regressions are with respect to the 100-year
252 climatological averages from the *piControl* runs). The intercepts of the regressions of $\overline{\delta R'_{\text{all}}}(\phi)$
253 versus global-mean surface-air temperature at each latitude are defined as the ‘adjusted’ radiative
254 forcings, $F(\phi)$ (Andrews et al. 2012). We use 100 years of simulation data following CO₂ qua-
255 drupling in these regressions and note that this method includes rapid climate adjustments as part
256 of the climate feedback. In the next section, we will compare the influences of rapid circulation
257 adjustments versus slower, temperature-mediated adjustments on the radiative response.

258 3. Atmospheric dynamics feedback in coupled climate models

259 a. Feedbacks versus latitude

260 The multimodel-mean all-sky and cloud feedbacks in the *abrupt4xCO2* simulations are plotted
261 in Figure 2, along with the thermodynamic, dynamic, and nonlinear components of these feed-

262 backs. The nonlinear component is substantially smaller than the dynamic component so for sim-
263 plicity we will mostly show the sum of the dynamic and nonlinear components rather than each
264 individually (a discussion of the factors controlling the magnitude of the nonlinear component is
265 provided later in this section).

266 1) ALL-SKY AND CLOUD FEEDBACKS

267 As expected, the total all-sky feedback is negative (i.e., decreasing net downward radiation at
268 TOA as climate warms) at all latitudes (Fig. 2a). The thermodynamic component typically dom-
269 inates the all-sky feedback but the dynamic component is substantial in the tropics, where its
270 magnitude is up to half of the magnitude of the thermodynamics feedback. Changes in large-scale
271 circulation produce a positive dynamics feedback close to the equator (the opposite of a negative
272 iris feedback) and contribute $+0.5 \text{ W m}^{-2}$ per degree global warming to the TOA imbalance in
273 that region. The positive all-sky dynamics feedback in the deep tropics is a result of cloud (Fig.
274 2b) and clear-sky feedbacks (not shown), and extends into the Southern Hemisphere subtropics,
275 though the magnitude is smaller than in the deep tropics. In the Northern Hemisphere subtropics,
276 the dynamics feedback is more complex, with a positive-negative dipole between 10°N and 30°N
277 (Fig. 2a).

278 The latitudinal structure of the dynamics feedback broadly reflects changes in mid-tropospheric
279 vertical velocity (Fig. 3): A strengthening of ascent in the core of the ITCZ under global warm-
280 ing intensifies moisture convergence in that region, which increases atmospheric water vapor and
281 enhances the longwave greenhouse effect (Fig. 4a,b). These positive longwave all-sky and cloud
282 feedbacks in the ITCZ are caused by strengthening ascent and moisture convergence in the middle
283 of the ITCZ with warming (Byrne and Schneider 2016b) and an upward shift in the altitude of high
284 clouds (Zelinka et al. 2012), which together produce a large and positive dynamics feedback on

285 the equator (Fig. 2a). The extent to which this dynamical longwave cloud feedback is consistent
286 with the fixed anvil temperature (FAT) hypothesis of Hartmann and Larson (2002) is a topic for
287 future work. The all-sky longwave feedback in the core of the ITCZ is counteracted somewhat
288 by a smaller negative shortwave feedback (Fig. 4a). This shortwave feedback is predominantly
289 associated with cloud processes (Fig. 4b) and points to increases in cloud fraction, thickness, and
290 albedo in the core of the ITCZ because of the strengthening and narrowing of the ascent in that
291 region; this pattern of tropical circulation changes has been termed the “deep-tropics squeeze”
292 (Lau and Kim 2015).

293 Moving from the core of the ITCZ to its equatorward flanks and into the subtropics, the long-
294 wave and shortwave components of the dynamics feedback change signs and tend to largely cancel
295 one another (Fig. 4). The weakening of ascending motion on the equatorward flanks of the cli-
296 matological ITCZ following CO₂ quadrupling (Fig. 3) causes reduced moisture convergence in
297 those regions and negative longwave clear-sky and cloud feedbacks; changes in free-tropospheric
298 clouds associated with a narrowing of the ITCZ are also likely to contribute to this negative long-
299 wave feedback. The longwave feedbacks are mostly cancelled by positive shortwave feedbacks at
300 the edges of the ITCZ and in the subtropics, and these shortwave components are also dominated
301 by cloud feedbacks (Fig. 4b). The positive shortwave cloud feedback in this region reflects a
302 decrease in the amount of shallow marine cloud with global warming (Bony and Dufresne 2005),
303 though the physical processes driving these changes in low-cloud amount are neither well under-
304 stood nor well represented in global climate models (Ceppi et al. 2017a). Because of the strong
305 cancellation between longwave and shortwave feedbacks on the equatorward flanks of the ITCZ
306 and in the subtropics, the overall all-sky dynamics feedback in this region is smaller than in the
307 core of the ITCZ (Fig. 4a).

308 2) FAST VERSUS SLOW DYNAMIC RESPONSES

309 Cloud-radiative feedbacks are typically divided into ‘cloud adjustments’, which emerge rapidly
310 in response to a radiative forcing, and surface temperature-mediated changes, which evolve slowly
311 as the climate warms (Andrews et al. 2012; Kamae et al. 2015). The dynamic components of
312 $\overline{\delta R'_{\text{all}}}(\phi)$ and $\overline{\delta R_{\text{cloud}}}(\phi)$ also have fast and slow components (Fig. 5a,c). Poleward of approxi-
313 mately 30° , the dynamic components are very similar in the first year (“fast”) and averaged over
314 years 91-100 (“slow”) following the quadrupling of CO_2 . Equatorward of 30° , however, there are
315 considerable differences between the fast and slow dynamic responses. In the deep tropics, close to
316 the equator, both the all-sky (Fig. 5a) and cloud-radiative dynamic responses (Fig. 5c) are strongly
317 positive (warming) 91-100 years after CO_2 quadrupling but are much smaller and negative in the
318 first year. This slow response is consistent with the ascent at the core of the ITCZ intensifying con-
319 siderably between the two time periods analyzed (not shown). The thermodynamic components
320 of $\overline{\delta R'_{\text{all}}}(\phi)$ and $\overline{\delta R_{\text{cloud}}}(\phi)$ emerge much more slowly (Fig. 5b,d), as expected by definition of the
321 thermodynamic components as being controlled by temperature changes.

322 The vertically-integrated atmospheric moist static energy (MSE) budget can be used diagnosti-
323 cally to understand the physical processes influencing tropical vertical velocity (Byrne and Schnei-
324 der 2016a). Changes in the net energy input to the atmosphere (e.g., through TOA radiative
325 anomalies) influence the MSE budget and consequently vertical velocity, and can be a fast re-
326 sponse to CO_2 forcing. On the other hand, poleward transport of MSE by large-scale mean ad-
327 vection and transient eddies also affects tropical vertical velocity and depends fundamentally on
328 meridional MSE gradients (Byrne and Schneider 2016a); these gradients would be expected to
329 evolve slowly in response to CO_2 forcing as sea-surface temperatures increase. It is unclear why
330 the dynamic components of changes in cloud-radiative effect [$\overline{\delta R_{\text{cloud}}}(\phi)$] and all-sky radiative

331 response $[\overline{\delta R'_{\text{all}}}(\phi)]$ evolve on different timescales at low and high latitudes [Ceppi et al. (2017b)
332 find a similar latitudinal dependence of timescales] – this is a topic for future work. The different
333 circulation-response timescales at different latitudes will be shown later to influence the global
334 feedback parameter and the response of the Earth system to CO₂ forcing.

335 Finally, there is only a weak correlation across CMIP5 models between the globally-averaged
336 fast and slow dynamic components of the all-sky radiative response ($r = 0.38$). This weak re-
337 lationship casts doubt on the feasibility of using the fast component to understand the long-term
338 influence of changes in atmospheric circulation on TOA radiation. This contrasts with the fast and
339 slow responses of tropical precipitation changes, which have been shown to be strongly determined
340 by the fast circulation response (Bony et al. 2013).

341 3) “AREA” AND “MIX” COMPONENTS OF THE DYNAMICS FEEDBACK

342 In an influential paper on the energetics of tropical climate, Pierrehumbert (1995) used a simple
343 two-box model to argue that “the warm pool sea surface temperature is very sensitive to [...] the
344 relative area of the dry versus convective regions”. Expressed in an equivalent way, Pierrehumbert
345 (1995) was suggesting that the ratio of the areas of tropical ascent versus descent regions is an
346 important controller of tropical climate. A widening of the Hadley circulation and a narrowing of
347 the ITCZ are expected to occur as climate warms, resulting in a decrease in the area of tropical
348 ascent relative to the area of tropical descent (Byrne and Schneider 2016b). We have calculated
349 the total dynamics feedback, but it is also of interest to isolate the portion of this feedback that is
350 due to a change in the up-down area ratio (e.g., due to ITCZ narrowing) and the portion that is due
351 to strengthening/weakening of the circulation within the distinct ascent and descent regimes (e.g.,
352 due to a slowdown of the Walker circulation).

353 We decompose each of the all-sky and cloud dynamics feedbacks, $\lambda_{\text{all}}^{\text{dy}}(\phi)$ and $\lambda_{\text{cloud}}^{\text{dy}}(\phi)$, into two
 354 parts: The first part, the ‘‘area component’’ $\lambda_{\text{area}}^{\text{dy}}(\phi)$, corresponds to changes in the up-down area
 355 ratio and the second part, the ‘‘mix component’’ $\lambda_{\text{mix}}^{\text{dy}}(\phi)$, corresponds to changes in the vertical-
 356 velocity distribution within the ascent and descent regimes. The two components of the all-sky
 357 dynamics feedback are expressed mathematically as follows:

$$\begin{aligned}
 \lambda_{\text{all}}^{\text{dy}}(\phi) &= \lambda_{\text{area}}^{\text{dy}}(\phi) + \lambda_{\text{mix}}^{\text{dy}}(\phi) \\
 &= \underbrace{\left(\int_{-\infty}^0 [R_{\text{all}}]^{\text{u}} [\delta A]^{\text{u}} d\omega + \int_0^{\infty} [R_{\text{all}}]^{\text{d}} [\delta A]^{\text{d}} d\omega \right)}_{\text{area component}} / [\delta T]^g + \\
 &\quad \underbrace{\left(\int_{-\infty}^0 (R_{\text{all}} - [R_{\text{all}}]^{\text{u}}) (\delta A - [\delta A]^{\text{u}}) d\omega + \int_0^{\infty} (R_{\text{all}} - [R_{\text{all}}]^{\text{d}}) (\delta A - [\delta A]^{\text{d}}) d\omega \right)}_{\text{mix component}} / [\delta T]^g
 \end{aligned} \tag{8}$$

358 where $[\cdot]^{\text{u}}$ indicates an average over all ascent regimes, $[\cdot]^{\text{d}}$ is an average over all descent regimes,
 359 and $[\delta T]^g$ is the global-mean surface-air temperature change as before. The dynamic component
 360 of the cloud feedback is decomposed into area and mix components in a similar way. Applying the
 361 decomposition (8) to the dynamics feedback (and estimating the area and mix components using
 362 least-squares regression, as discussed above), we find that the area and mix components both
 363 contribute to the all-sky and cloud dynamics feedbacks (Fig. 6). For the all-sky feedback, the area
 364 and mix components are both positive in the deep tropics (Fig. 6a) and contribute approximately
 365 equally to the strong positive dynamics feedback in that region. These positive contributions
 366 to the dynamics feedback in the deep tropics are a result of a strengthening of upward motion
 367 within ascending regions (mix component) and small increases in the ascending area within the
 368 ITCZ (area component). On the equatorward flanks of the ITCZ, the area component is a weak
 369 negative feedback reflecting atmospheric drying and reduced shortwave absorption and longwave
 370 greenhouse effect due to the narrowing of the ITCZ. Both the area and mix components decay

371 in magnitude moving away from the equator. Interestingly, the mix component dominates the
372 dynamic component of the cloud feedback (Fig. 6b).

373 Overall, the influence of changes in the tropical up-down area ratio on climate feedbacks is weak
374 compared to the influence of thermodynamic processes (cf. Fig. 2). This result contrasts with the
375 findings of Pierrehumbert (1995) based on a simple two-box model of the tropical climate, and
376 casts doubt on the ability of changes in the widths of the ITCZ or Hadley circulation to strongly
377 influence global climates of the past or the future.

378 4) SIZE OF THE NONLINEAR COMPONENT

379 We found that the nonlinear components of the radiative feedbacks are small relative to the dy-
380 namic and thermodynamic components (Fig. 2). However, it is interesting to consider the factors
381 controlling the magnitudes of these nonlinear terms. To gain insight into the nonlinear components
382 of the all-sky and cloud feedbacks [(6) and (7)], it is instructive to assume the relationship between
383 TOA radiation and vertical velocity is linear, i.e., $R(\omega) = a + b\omega$ (here for convenience R repre-
384 sents either R_{all} or R_{cloud}). The assumption of linearity is reasonable in climate models over the
385 majority of tropical vertical-velocity regimes for both the cloud-radiative effect (Fig. 1a) and clear-
386 sky fluxes (Fig. 1b). Taking the definitions of the dynamic and nonlinear components of the all-
387 sky radiative response (4) or changes in cloud-radiative effect (5), substituting in $R(\omega) = a + b\omega$
388 and rearranging, the ratio of the nonlinear component to the dynamic component can be shown
389 to be equal to $\delta b/b$, where δb in this case denotes the change in the slope of the linear fit to
390 $R(\omega)$ between the *piControl* and *abrupt4xCO2* runs. For the cloud-radiative effect, $\delta b/b \approx -20\%$
391 meaning that the sensitivity of the cloud-radiative effect to changes in mid-tropospheric vertical
392 velocity decreases by 20% under global warming in the simulations analyzed here. This weaken-
393 ing of the relationship between vertical velocity and cloud-radiative effect as the climate warms

394 could be a result of cloud-masking effects associated with rising CO₂ and water-vapor concentra-
 395 tions in the atmosphere, and it also suggests that mid-tropospheric vertical velocity alone is not a
 396 good predictor of cloud changes. Our simple analysis suggests that the nonlinear component of
 397 changes in cloud-radiative effect (or the nonlinear component of the all-sky radiative response)
 398 is approximately 20% of the magnitude of the dynamic component, which roughly matches the
 399 nonlinear feedbacks calculated explicitly using the decompositions [(6) and (7)] (cf. Figure 2).

400 *b. Global atmospheric dynamics feedback*

401 Averaged over all latitudes, TOA radiative anomalies induced by CO₂ quadrupling are mostly a
 402 result of thermodynamic processes (see the results for one representative climate model in Figure
 403 7). Although the atmospheric dynamics feedback is relatively important at low latitudes (Fig. 2a),
 404 there is a large degree of cancellation between positive and negative dynamics feedbacks when the
 405 global average is calculated. In what follows, we explain why the atmospheric dynamics feedback
 406 is small when averaged over large regions.

407 1) MASS BUDGET CONSTRAINT

408 That the global atmospheric dynamics feedback is small is perhaps not surprising when the con-
 409 straint of the atmospheric mass budget is taken into account. Within a region that has a negligible
 410 net atmospheric mass flux across its lateral boundaries, such as the globe or the region occupied
 411 by the Hadley circulation, the steady-state atmospheric mass budget can be expressed as:

$$\int_{-\infty}^0 \omega A(\omega) d\omega = - \int_0^{\infty} \omega A(\omega) d\omega, \quad (9)$$

412 where $A(\omega)$ here is the normalized area PDF of vertical velocity defined for the specified region
 413 (e.g., the globe). Expression (9) stipulates that in the steady-state atmosphere, the area of up-
 414 ward motion multiplied by the average upward velocity is equal to the area of downward motion

415 multiplied by the average downward velocity. We now illustrate how this mass budget places a
416 strong constraint on the ability of changes in the atmospheric circulation to influence global TOA
417 radiation and climate.

418 As in our discussion of the nonlinear component above, a useful starting point is to assume the
419 dependence of TOA radiation (either the all-sky radiative response or the cloud-radiative effect) on
420 mid-tropospheric vertical velocity is linear, i.e., $R(\omega) = a + b\omega$. Linearity is not an unreasonable
421 assumption for climate models, at least for tropical vertical-velocity regimes with a heavy area
422 weighting (Fig. 1). The linearity assumption is also supported by satellite observations of the
423 cloud-radiative effect (Wyant et al. 2006). Using this assumption, the dynamic components of (4)
424 or (5) are given by: $\int_{-\infty}^{\infty} R(\omega)\delta A(\omega)d\omega = a \int_{-\infty}^{\infty} \delta A(\omega)d\omega + b \int_{-\infty}^{\infty} \omega\delta A(\omega)d\omega$. By construction
425 of normalized area PDFs, $\int_{-\infty}^{\infty} \delta A(\omega)d\omega = 0$, and from the atmospheric mass budget (9) it follows
426 that $\int_{-\infty}^{\infty} \omega\delta A(\omega)d\omega = 0$. Consequently, for a strictly linear relationship between TOA radiation
427 and vertical velocity, the impact of changes in atmospheric circulation on TOA radiative anomalies
428 averaged over a closed-mass region is expected to be zero. A slightly modified argument can be
429 used to show that the nonlinear components of (4) and (5) are also zero for linear $R(\omega)$. The
430 implications of a linear $R(\omega)$ for the dynamic component of changes in cloud-radiative effect have
431 been discussed previously by Wyant et al. (2006).

432 2) EFFECT ON GLOBAL CLIMATE CHANGE

433 Although small in magnitude, the globally-averaged atmospheric dynamics feedback neverthe-
434 less has an influence on the climate response to CO₂ quadrupling (Figs. 7 and 8). For the equilib-
435 rium climate response (ECR), defined here as the equilibrium change in near-surface temperature
436 following CO₂ quadrupling calculated using the regression method of Gregory et al. (2004), the
437 dynamics feedback marginally increases the surface warming (Figs. 7 and 8a) – the multimodel-

438 mean ECR is approximately 0.5 K larger than that predicted by the thermodynamics feedback
439 alone (0.5 K corresponds to 8% of the multimodel-mean ECR in these simulations). Interestingly,
440 the global dynamics feedback is robustly *positive* across CMIP5 models (Fig. 8). This contrasts
441 with the hypothesis that tropical circulation changes may be a *negative* iris-type feedback on cli-
442 mate change (e.g., Lindzen et al. 2001). Despite being a small and positive feedback on climate
443 change, there is no correlation between the all-sky dynamics feedback (globally averaged) and
444 ECR, indicating that differences in atmospheric circulation responses across models cannot be
445 used to explain intermodel spread in ECR. In contrast, there is a strong correlation of $r = 0.83$
446 between the thermodynamics feedbacks and ECRs across models.

447 The global feedback parameter, or the slope of the linear regression of the all-sky radiative
448 response versus near-surface temperature anomalies, describes the efficiency with which the Earth
449 system equilibrates a radiative forcing. The feedback parameter is the sum of all the individual
450 feedback processes (e.g., water vapor, cloud, lapse rate) and is typically negative, meaning that
451 the net downward flux of radiation at TOA decreases as the surface temperature increases during
452 equilibration. The atmospheric dynamics feedback robustly makes the global feedback parameter
453 less negative (Figs. 7 and 8b), implying that the Earth system is less efficient at eliminating a TOA
454 forcing because of circulation changes. This reduction in the magnitude of the feedback parameter
455 is consistent with the all-sky radiative response associated with circulation changes becoming less
456 negative over time following CO₂ quadrupling (Figs. 5a and 7).

457 The climate models examined here are unable to simulate processes such as convective aggrega-
458 tion that have been put forward as potentially important influencers of global climate (Bony et al.
459 2016). Furthermore, climate models struggle to accurately simulate clouds (Pincus et al. 2008),
460 and it is plausible that important nonlinearities in the $R_{\text{all}}(\omega, \phi)$ and $R_{\text{cloud}}(\omega, \phi)$ relationships, not
461 captured by climate models, could lead to dynamics feedbacks that are substantially non-zero on

462 global scales. In the next section we will use an idealized GCM with prescribed TOA forcings
463 to investigate how large atmospheric dynamics feedbacks, potentially induced by the effects men-
464 tioned above but not simulated by the current generation of climate models, could influence global
465 climate.

466 **4. Effectiveness of high- versus low-latitude dynamics feedbacks**

467 *a. Idealized forcing simulations*

468 We use a moist idealized GCM perturbed by stylized TOA longwave forcing profiles to in-
469 vestigate in a general way how the latitudinal structure of the atmospheric dynamics feedback
470 could impact global climate. Simulations are performed using the slab-ocean aquaplanet GCM
471 of O’Gorman and Schneider (2008), which is based upon the model of Frierson et al. (2006) and
472 Frierson (2007). The GCM is run at a horizontal spectral resolution of T42 with 10 vertical levels
473 and the surface heat capacity is equivalent to 1m of liquid water. The model uses a simplified two-
474 stream gray radiation scheme with a prescribed longwave optical thickness; shortwave fluxes are
475 specified as a function of latitude and longitude [see O’Gorman and Schneider (2008) for full de-
476 tails of the radiation scheme]. It is important to note that this model does not simulate water-vapor
477 or cloud feedbacks.

478 We perform a control simulation (without TOA longwave forcing) that has a climate roughly
479 similar to that of the present-day Earth, along with three perturbed simulations with prescribed
480 longwave forcings (Fig. 9a): A tropical forcing from $\phi = -30^\circ$ to 30° and an extratropical forcing
481 ($\phi = -90^\circ$ to -30° and $\phi = 30^\circ$ to 90°), both with a global-mean forcing of 4 W m^{-2} , and a
482 global wave forcing with a maximum amplitude of 4 W m^{-2} and a global mean of 0 W m^{-2} . Each
483 forcing has a $\cos 3\phi$ dependence on latitude and is invariant with longitude. To prescribe the TOA

484 longwave forcings, we simply impose longwave radiative fluxes (positive downward) at the top
485 level of the model with patterns as shown in Figure 9a. Each simulation is spun up for 1000
486 days with averages taken over the subsequent 2000 days. The magnitudes of the stylized forcings
487 are considerably larger than what we have calculated for the dynamic component of the all-sky
488 radiative response in the fully-coupled simulations (cf. Fig. 5), and the latitudinal structures
489 imposed here are also highly idealized. However, our objective is to understand the effectiveness
490 of high-latitude versus low-latitude TOA anomalies (due to changes in atmospheric circulation) in
491 changing global climate, and for this purpose using an idealized GCM and stylized forcings is an
492 appropriate first step.

493 *b. Surface temperature responses*

494 Although the tropical and extratropical forcings have equal magnitudes in the global mean, they
495 induce strikingly different surface temperature responses (Fig. 9b). The global temperature re-
496 sponse to extratropical forcing (1 K) is greater than twice the magnitude of the response to tropical
497 forcing (0.4 K). This stronger sensitivity of climate to extratropical versus tropical forcing has
498 also been shown for surface forcings (Rose et al. 2014; Rugenstein et al. 2016) – ocean heat
499 uptake at high latitudes cools the climate more efficiently than the same heat uptake at low lati-
500 tudes. The difference in temperature responses between low- and high-latitude forcings (a factor
501 of approximately 2) found in our idealized simulations is somewhat smaller than for full-physics
502 simulations forced by ocean heat uptake patterns (a factor of approximately 3–4; Rose et al. 2014).
503 The enhanced contrast in global temperature responses for high- versus low-latitude forcings in
504 full-physics simulations (Rose et al. 2014) relative to the idealized simulations is due to the full-
505 physics simulations including water-vapor and cloud feedbacks, which further strengthen differ-
506 ences in the forcing responses. To further emphasize how the latitudinal pattern of TOA forcing

507 influences global temperature, the global-wave forcing has a global-mean magnitude of zero (Fig.
508 9a) and yet induces a surface-temperature cooling of 0.2 K. That the global-mean temperature can
509 change in response to a forcing with latitudinal structure but zero global mean has been shown in
510 previous studies (e.g., Rose and Ferreira 2013; Rose and Rencurrel 2016).

511 The physical mechanisms leading to the enhanced global climate response to high-latitude ver-
512 sus low-latitude forcings have received considerable attention (e.g., Hansen et al. 1997; Langen
513 and Alexeev 2007; Lu and Cai 2010; Pithan and Mauritsen 2014; Rose et al. 2014; Roe et al. 2015).
514 Fundamentally, the contrasting temperature responses are driven by different sensitivities of pole-
515 ward atmospheric energy transport to high-latitude and low-latitude temperature perturbations, as
516 well as latitudinal variations in local radiative feedbacks [see Rose et al. (2014) for an insightful
517 discussion on this topic]. Consequently, although the idealized GCM used here omits a variety
518 of processes which might be expected to change the magnitude of the climate response to TOA
519 forcing (e.g., water-vapor and ice-albedo feedbacks), we believe the larger climate sensitivity to
520 high-latitude forcing is a robust result that is grounded in well-established physical mechanisms. It
521 should be noted that despite global surface temperature being more sensitive to a 1 W m^{-2} forcing
522 prescribed at high latitudes versus at low latitudes, the equilibrium climate sensitivity in CMIP5
523 models and its inter-model spread is dominated by tropical forcings and feedbacks because the
524 tropics makes up a large fraction of Earth's total area and because cloud feedbacks are more un-
525 certain across models in the tropics than at higher latitudes (Vial et al. 2013).

526 The idealized simulations discussed in this section have important implications for the atmo-
527 spheric dynamics feedback. Although the influence of circulation changes on Earth's global radia-
528 tive balance is constrained to be weak as discussed above, the effect of circulation changes on the
529 TOA radiative response at different latitudes is non-negligible (Fig. 2a), and the global temperature
530 response to circulation changes is likely to depend on this latitudinal structure. Furthermore, dis-

531 cussions of how the atmospheric circulation response might mitigate climate change have focused
532 largely on tropical iris-type mechanisms (e.g., Pierrehumbert 1995; Lindzen et al. 2001; Mauritsen
533 and Stevens 2015; Bony et al. 2016). However, based on the idealized simulations presented here,
534 a 10 W m^{-2} tropical TOA forcing would be needed to change global temperature by 1 K (a more
535 realistic model with water-vapor and cloud feedbacks would be expected to require a smaller TOA
536 forcing to change global temperature by 1 K, as discussed above). Our analysis of the influence of
537 atmospheric dynamics on TOA radiation in CMIP5 models suggests a circulation-induced tropical
538 TOA radiative anomaly of no greater than $1\text{--}2 \text{ W m}^{-2}$ under CO_2 quadrupling (Fig. 5a), which
539 would have only a minor influence on global temperature. Furthermore, an observational analy-
540 sis shows that convective aggregation also has only a weak influence on TOA radiation [of order
541 1 W m^{-2} (Tobin et al. 2012)]. Consequently, we argue that the relative insensitivity of global tem-
542 perature to tropical TOA perturbations, combined with these perturbations due to changes in the
543 atmospheric circulation being constrained to be small on large scales, means that tropical iris-type
544 mechanisms are unlikely to exert a substantial influence on past or future climate change.

545 **5. Summary and discussion**

546 The response of the atmospheric circulation to radiative forcing, via its influence on clouds and
547 water vapor, has been put forward as a potentially important negative feedback on climate change
548 (Lindzen et al. 2001). Here, using a method based on that of Bony et al. (2004), we have calcu-
549 lated for the first time the atmospheric dynamics feedback at each latitude in a range of climate
550 models subjected to an abrupt CO_2 quadrupling. For the all-sky feedback, the dynamic compo-
551 nent is generally smaller than the thermodynamic component, though is not entirely negligible,
552 particularly in the tropics, where it is positive with a magnitude of approximately $0.5 \text{ W m}^{-2} \text{ K}^{-1}$.
553 For the cloud feedback, the dynamic component is comparable to the thermodynamic component

554 close to the equator. The latitudinal pattern of the dynamics feedback reflects previously iden-
555 tified changes in the tropical circulation under global warming including enhanced ascent in the
556 core of the ITCZ and a narrowing of this convergence zone. This coupling between radiation and
557 circulation emphasizes the necessity of understanding the sensitivity of large-scale atmospheric
558 dynamics to climate change in order to fully understand climate feedbacks at low latitudes.

559 At individual latitudes in the tropics, the atmospheric dynamics feedback is comparable to the
560 thermodynamics feedback. However, a large degree of cancellation between positive and negative
561 dynamics feedbacks at different latitudes results in the thermodynamic component dominating
562 the global climate response. That the global-mean atmospheric dynamics feedback is small is
563 shown to be a consequence of (i) the steady-state atmospheric mass budget and (ii) approximately
564 linear relationships between mid-tropospheric vertical velocity and the cloud-radiative effect and
565 all-sky fluxes, though these relationships between vertical velocity and radiation require further
566 investigation in high-resolution simulations that permit clouds and explicitly simulate convection.
567 This atmospheric-mass constraint on the dynamics feedback raises doubts regarding the ability of
568 iris-type mechanisms (e.g., convective aggregation) to limit future climate change.

569 Although the global radiative anomalies induced by circulation changes are small relative to
570 anomalies induced by thermodynamic processes (e.g., temperature-driven changes in water va-
571 por), they do affect the climate response to CO₂ quadrupling. Circulation changes are a positive
572 feedback on global climate change (in contrast to iris-type negative feedbacks) and make the Earth
573 system less efficient at equilibrating an imposed radiative forcing. Specifically, the dynamics feed-
574 back increases the multimodel-mean ECR by 0.5 K (or approximately 8% of the total warming)
575 relative to a calculation of ECR using the thermodynamics feedback alone.

576 We performed idealized simulations to assess how hypothetical circulation-induced TOA
577 anomalies at low and high latitudes, such as those due to unresolved dynamical processes includ-

578 ing convective aggregation, could affect global climate. We showed that a high-latitude forcing is
579 twice as effective as a low-latitude forcing in changing global surface temperature (interestingly
580 and in agreement with previous studies, we also found that a forcing with a global average of
581 zero can induce a non-zero temperature response). These results further strengthen our argument
582 that low-latitude dynamical responses to climate change – such as convective aggregation or ITCZ
583 narrowing – are unlikely to greatly influence global climate unless they are capable of inducing
584 very large TOA anomalies. On the other hand, circulation changes that create TOA anomalies at
585 higher latitudes (for example poleward shifts in the storm tracks or deflections of the mid-latitude
586 jet by large continental ice sheets in cold climates) are expected to be more than twice as effective
587 at changing global temperature and consequently should be a priority to understand and quantify.

588 Our findings suggest a number of avenues for future research. Decades of satellite measure-
589 ments now enable observational estimates of the atmospheric dynamics feedback; such estimates
590 would be a useful benchmark with which to assess our climate-model results. These observa-
591 tions could also be examined in detail to assess whether non-linearities in observed $R_{\text{all}}(\omega, \phi)$ and
592 $R_{\text{cloud}}(\omega, \phi)$ relationships are sufficiently strong to generate atmospheric dynamics feedbacks that
593 could strongly influence global climate. We have considered here the effect of changes in resolved
594 dynamical processes on climate in global models, but it would also be insightful to perform similar
595 analyses in high-resolution convection-permitting simulations to assess how small-scale dynami-
596 cal processes couple to TOA radiation. Our focus has been the influence of atmospheric circulation
597 changes on TOA radiation and global temperature. However, the global precipitation response to
598 climate change is linked to TOA as well as to surface radiation (e.g., O’Gorman et al. 2012). An
599 interesting topic for future work would be to understand the contribution of circulation changes
600 to the atmospheric (rather than to the TOA) radiation budget and consequently to the response of
601 global precipitation to climate change (Su et al. 2017).

602 *Acknowledgments.* We thank Paulo Ceppi, Reto Knutti, Paul O’Gorman, Brian Rose, Maria Ru-
603 genstein, and the Editor for helpful comments, suggestions, and discussions. M.P.B. acknowledges
604 support from the Imperial College London Research Fellowship Scheme.

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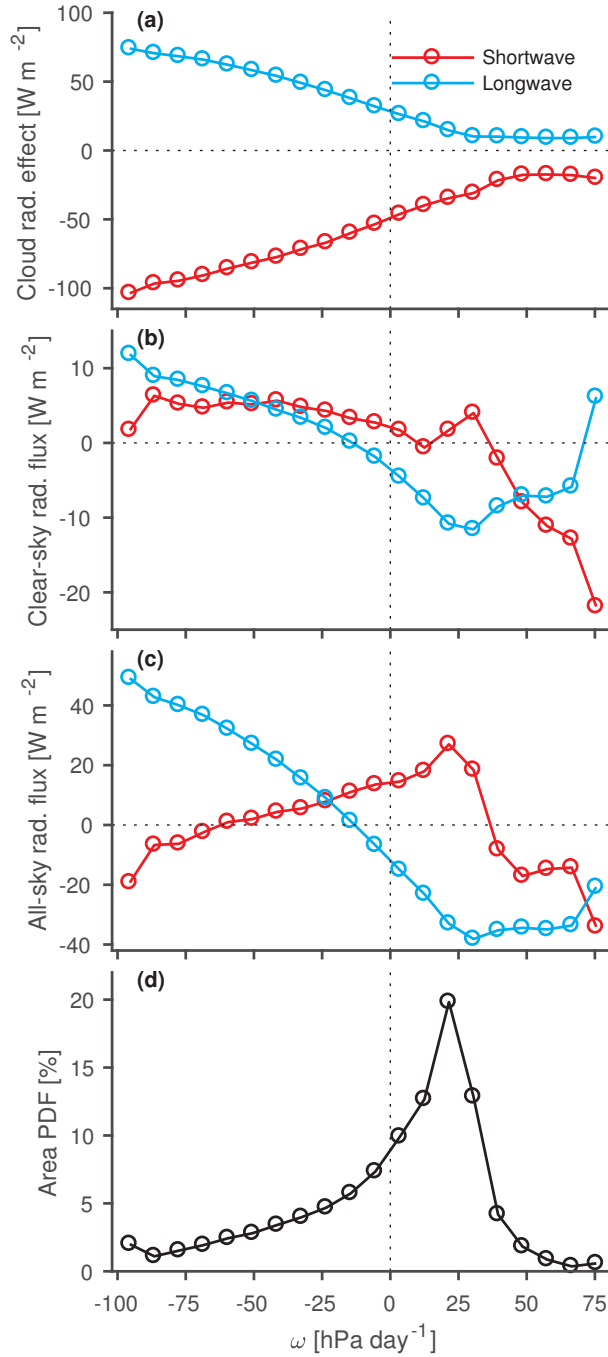
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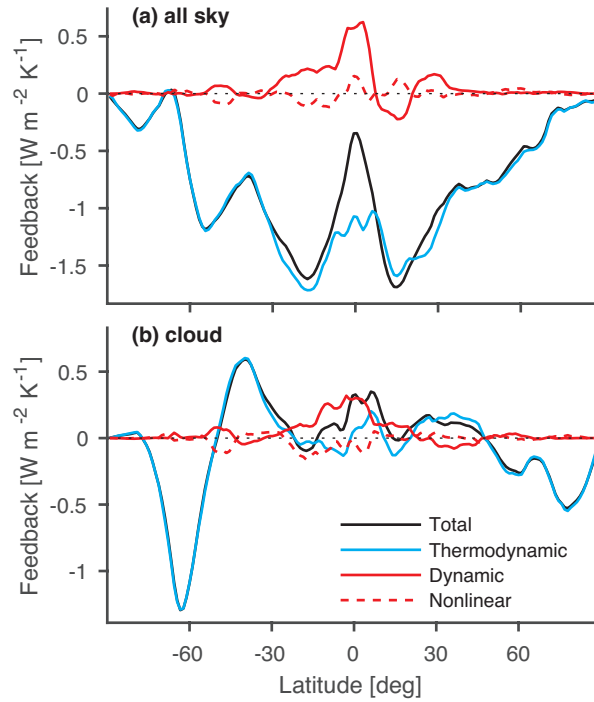
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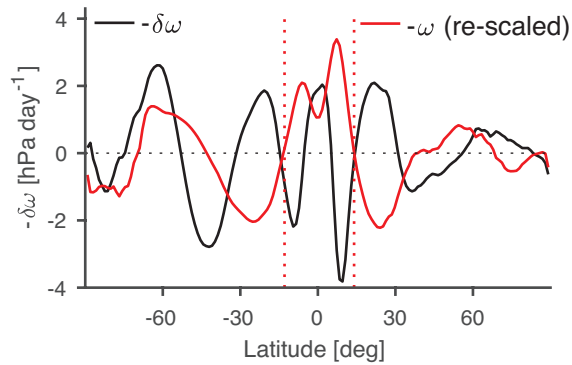
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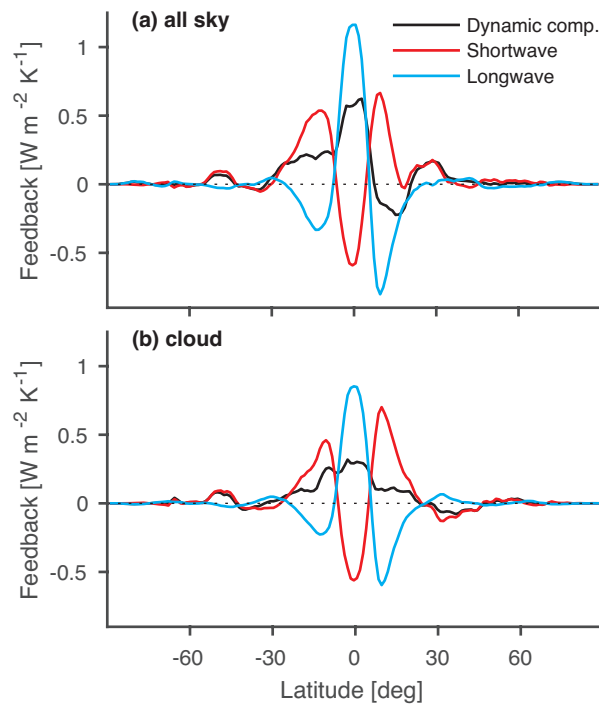
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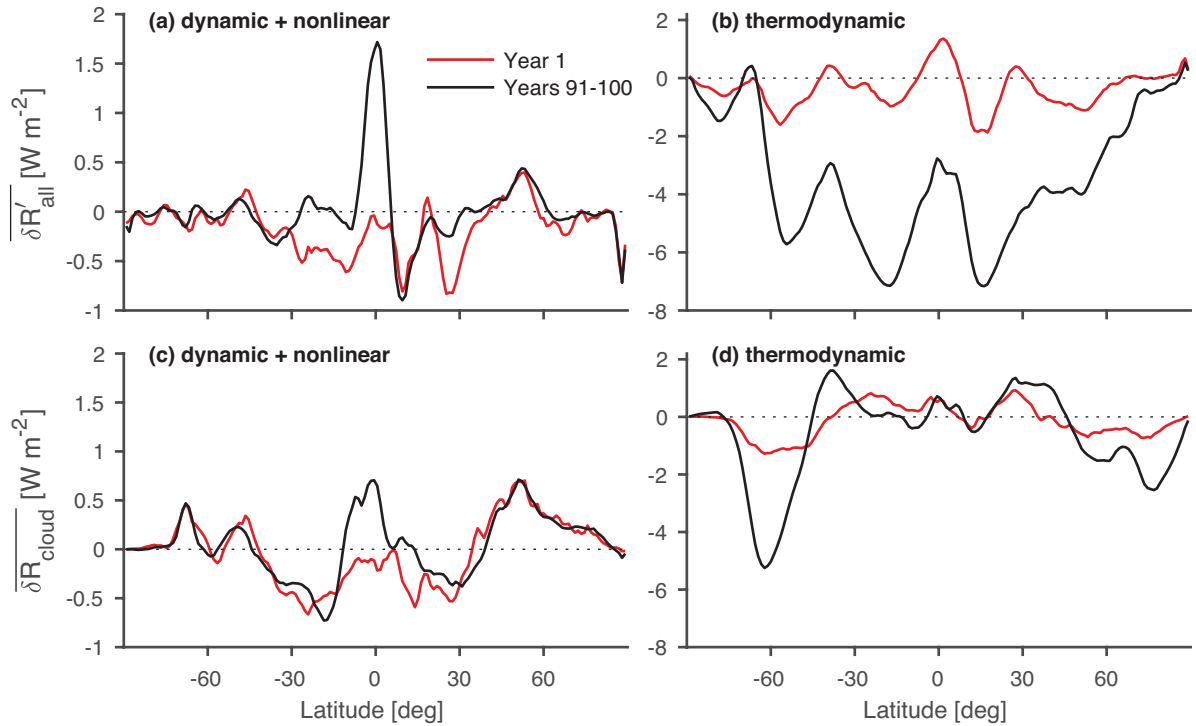
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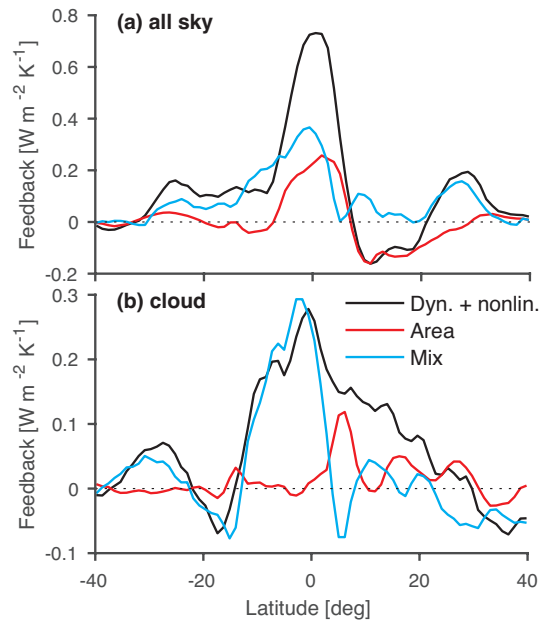
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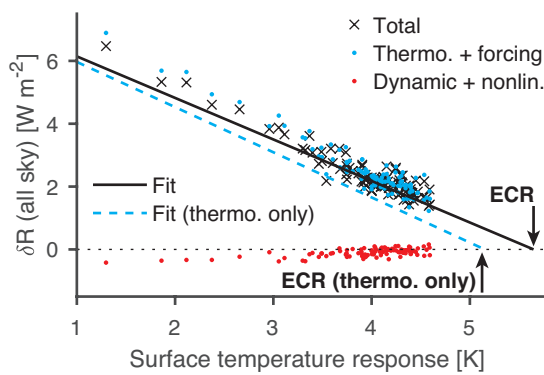
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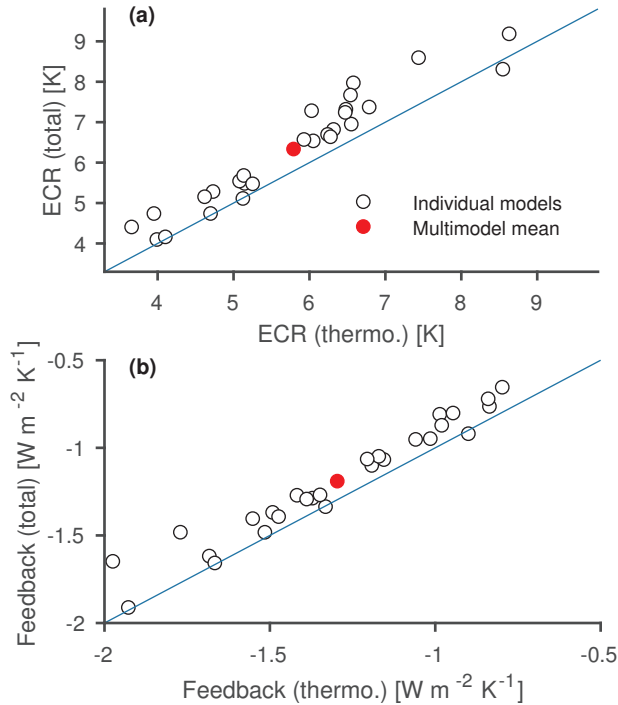
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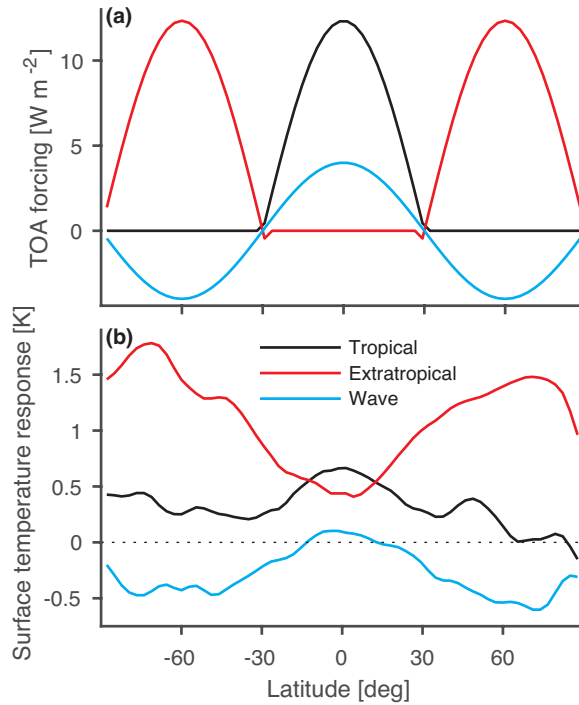
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